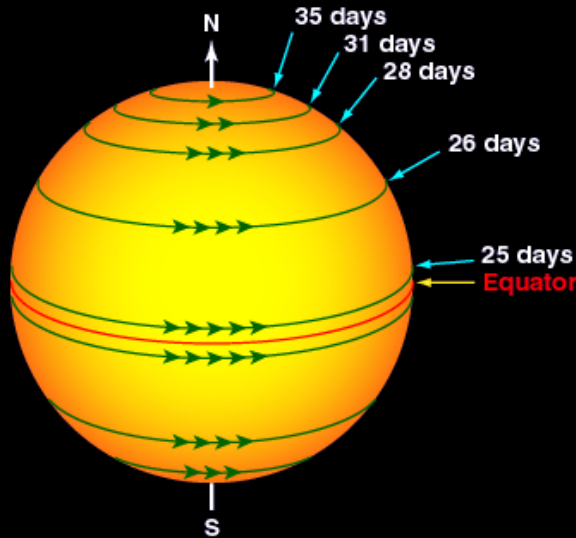
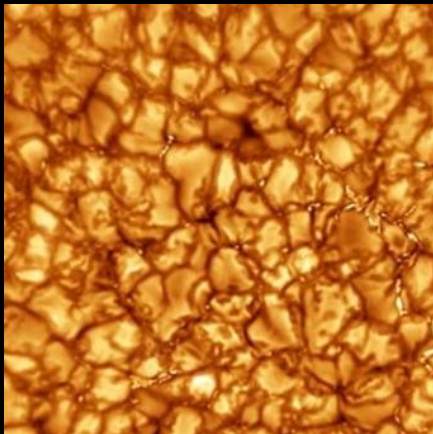


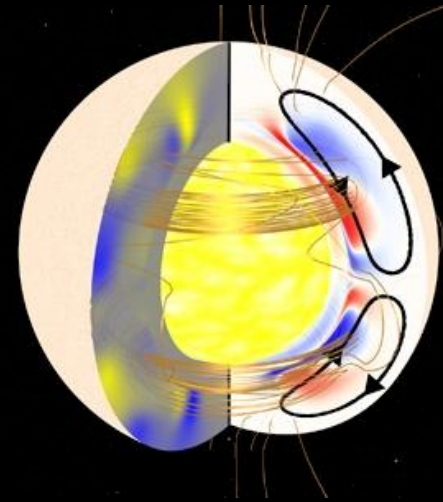
# Understanding the Solar Magnetic Cycle



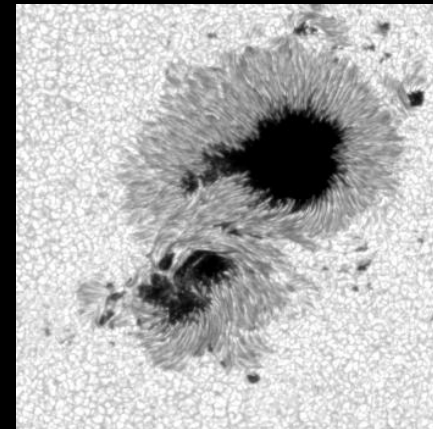
**Differential Rotation**



**Turbulent Convection**



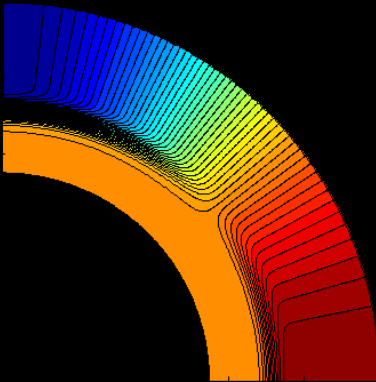
**Meridional Flow**



**Sunspot Eruption**

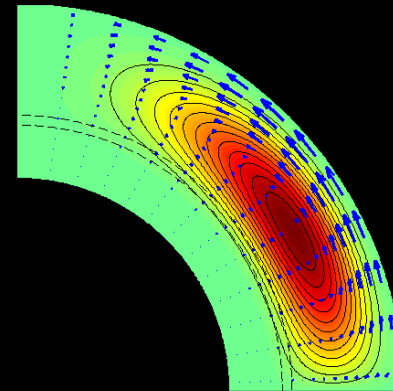
# Modeling the Solar Magnetic cycle

## Differential Rotation



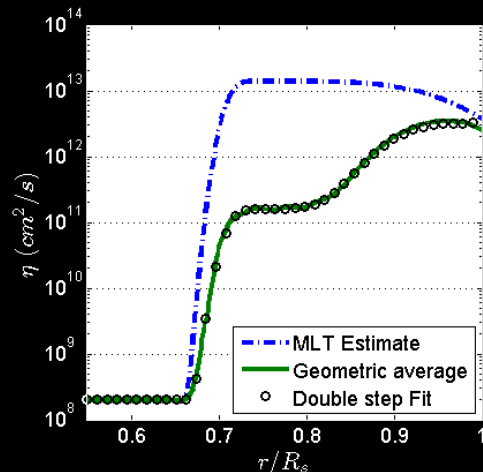
Charbonneau et al. 1999

## Meridional Flow



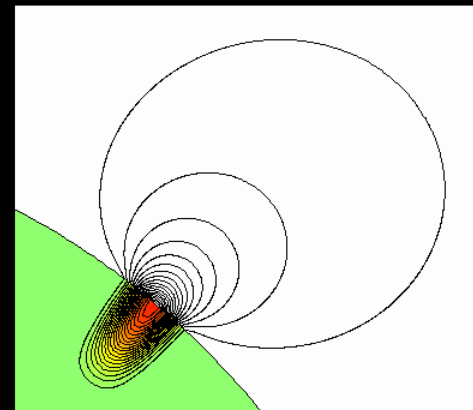
Muñoz-Jaramillo, Nandy  
& Martens 2009

## Turbulent Diffusivity



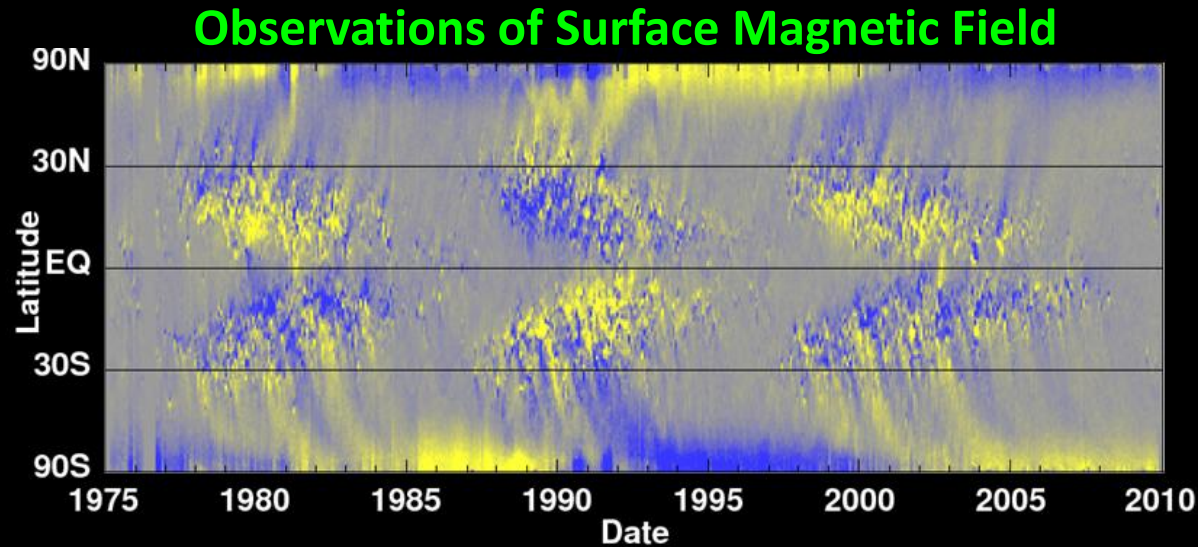
Muñoz-Jaramillo, Nandy  
& Martens 2011

## Sunspot Eruption

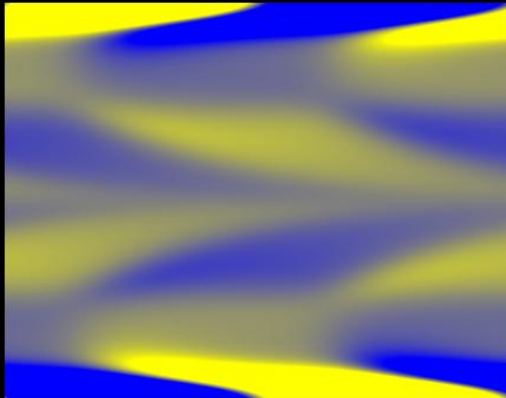


Muñoz-Jaramillo et al. 2010

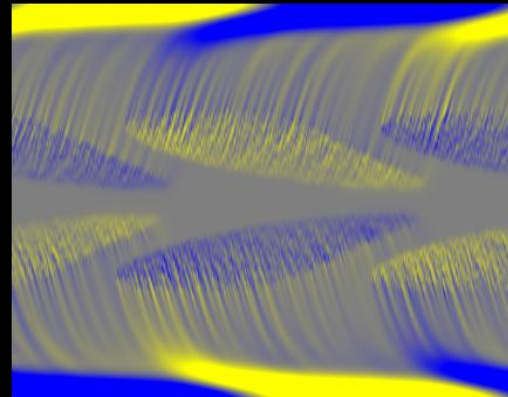
# How does our model compare with observations and other models?



**Traditional Models**



**Improved Model**



# Understanding the Extended Solar Minimum

## 1. Large amount of days without sunspots.



Our model includes a realistic algorithm for sunspot eruption.

## 2. Weak polar field strength.



Our model captures successfully the dynamics of polar field generation